

Spatially Homogeneous Universes Which Admit Source-Free Magnetic Fields.

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In recent years interest in the Einstein-Cartan theory of space-time⁽¹⁻³⁾ has grown significantly. In the theory the space-time manifold is endowed with a torsion (non-symmetric connection) field, which couples to the spin density of matter in the universe.

Cosmological models based on the Einstein-Cartan theory are usually studied under the assumption that, having aligned the spins of the particles along a particular direction, the magnetic field that might be present in the universe decouples from the magnetic dipoles of the matter content⁽⁴⁾.

Non coupling between matter and electromagnetic fields in the universe is also assumed in most cosmological models based on the theory of general relativity^(5,6).

It is of interest, therefore, to determine the class of spatially homogeneous space-times which admit the presence of source-free magnetic fields. This can be done simultaneously in the Einstein-Cartan theory and in general relativity, because it has been shown⁽²⁾, that the sourceless equations of Maxwell are the same in both theories, namely

$$(1) \quad dF = 0$$

and

$$(2) \quad d*F = 0,$$

where F and $*F$ are the electromagnetic field 2-form and its dual, respectively.

In a spatially homogeneous space-time manifold, *i.e.* one which is locally invariant under a group of isometries G_3 simply transitive on spacelike surfaces, we can choose a set of basic 1-forms Θ^a ($a = 0, 1, 2, 3$) with $\Theta^0 = dt$, such that the hypersurfaces of homogeneity are given by

$$(3) \quad t(x^a) = \text{constant},$$

(1) F. W. HEHL: *Gen. Rel. and Gravit.*, **4**, 333 (1973).

(2) F. W. HEHL: *Gen. Rel. and Gravit.*, **5**, 491 (1974).

(3) A. TRAUTMAN: *Ann. N. Y. Acad. Sci.*, **262**, 241 (1975).

(4) A. K. RAYCHAUDHURI: *Phys. Rev.*, **12**, 952 (1975).

(5) D. R. BRILL: *Phys. Rev. Ser. B*, **133**, 845 (1964).

(6) D. G. TSOUBELIS: Ph. D. Dissertation, New York, N. Y., 1976 (unpublished).

where the x^α 's are local co-ordinates. With respect to such a basis, we can write the magnetic-field 2-form as

$$(4) \quad F = \frac{1}{2} \varepsilon^i{}_{jk} B_i \Theta^j \times \Theta^k,$$

where $B_i = B_i(t)$, $\varepsilon^i{}_{jk}$ is the totally antisymmetric symbol ($\varepsilon^1{}_{23} = 1$), and Latin indices run from 1 to 3. The dual of F becomes

$$(5) \quad *F = B_i \Theta^0 \times \Theta^i.$$

In the above basis, on the other hand, we have (⁷)

$$(6) \quad d\Theta^\alpha = -\frac{1}{2} C_{\beta\gamma}{}^\alpha \Theta^\beta \times \Theta^\gamma,$$

where $C_{\beta\gamma}{}^\alpha = C_{\beta\gamma}{}^\alpha(t) = C_{[\beta\gamma]}{}^\alpha$.

Substituting (4) and (5) into (1) and (2), respectively, and using (6), we obtain the following set of equations:

$$(7) \quad \frac{dB^i}{dt} + 2C_{0j}{}^i B^j = 0,$$

$$(8) \quad B_i \varepsilon^l{}_{mn} C_{jk}{}^m \varepsilon^{jkn} = 0$$

and

$$(9) \quad \varepsilon^{ijk} C_{jk}{}^l B_l = 0.$$

One can now write the $C_{jk}{}^i$'s in terms of a symmetric relative tensor n^{ij} and a relative vector a_i as

$$(10) \quad C_{jk}{}^i = \varepsilon_{jkl} n^{li} + \delta_k^i a_j - \delta_j^i a_k,$$

where $n^{ij} = n^{ij}(t) = n^{(ij)}$, $a_i = a_i(t)$, and

$$(11) \quad n^{ij} a_j = 0.$$

When (10) is substituted in (8) and (9), they give

$$(12) \quad B^j a_j = 0$$

and

$$(13) \quad (n^{ij} - \varepsilon^{ijk} a_k) B_j = 0,$$

respectively.

Now, if $a_i = 0$ (class A space-times), it follows from (13) that $\det(n^{ij}) = 0$. In this case, when the rank of the matrix (n^{ij}) is equal to 2, 1 or zero, then the group types

(⁷) A. R. KING and G. F. R. ELLIS: *Commun. Math. Phys.*, **31**, 209 (1973).

is the Bianchi type VI₀ or VII₀, II and I, respectively. (See (7) for the Bianchi-Schücking-Behr classification of group types based on the solution of the characteristic equation of (n^{ij}) .)

If $a_i \neq 0$ (class B space-times), then it follows from (11) that $\det(n^{ij}) = 0$, again. This, along with (12) and (13), implies that

$$(14) \quad [n^{ij}n_{ij} - (n_k^k)^2]/2a^k a_k = +1.$$

(It can be easily proven that (14) holds, by using an orthonormal frame where $n^{ij} = \text{diag}(0, n^2, n^3)$, $a_i = (a, 0, 0)$, and $B^i = (0, B^2, B^3)$.) Thus, in this case the group type must be VI₋₁, or Bianchi type III.

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